

# **Bayesian Analysis for Complex Physical Systems Modeled by Computer Simulators: Current Status and Future Challenges**

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practical/methodological (how can we work out what climate is likely to be?)

and

foundational (why should our methods work and what do our answers mean?)

## Examples

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The aim is scientific - to gain information about the physical processes underlying the Universe.

**Climate change** Large scale climate simulators are constructed to assess likely effects of human intervention upon future climate behaviour.

Aims are both scientific - much is unknown about the large scale interactions which determine climate - and also very practical, as such simulators provide evidence for the importance of changing human behaviour before possibly irreversible changes are set into motion.

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- (ix) **decision uncertainty** (to use the model to influence real world outcomes, we need to relate things in the world that we can influence to inputs to the simulator and through outputs to actual impacts. These links are uncertain.)

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[In a climate model,  $y_h$  might correspond to historical climate outcomes over space and time,  $y$  to current and future climate, and the “decisions” might correspond to different policy relevant choices such as carbon emission scenarios.]

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If observations,  $z$ , are made without error and the model is perfect reproduction of the system, we can write  $z = f_h(x^*)$ , invert  $f_h$  to find  $x^*$ , learn about all future components of  $y = f(x^*)$  and choose decision elements of  $x^*$  to optimise properties of  $y$ .

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In practice, the observations  $z$  are made with error, and model is not the same as physical system so we must separate the uncertainty representation into two relations and carry out statistical inversion/optimisation:

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COMMENT And we still haven't accounted for condition uncertainty, multi-model uncertainty, etc.

## Current state of the art

Many people work on different aspects of these uncertainty analyses

Great resource: the Managing Uncertainty in Complex Models web-site

<http://www.mucm.ac.uk/> (for references, papers, toolkit, etc.)

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and it is hard!

## RAPID-WATCH

What are the implications of RAPID-WATCH observing system data and other recent observations for estimates of the risk due to rapid change in the MOC? In this context risk is taken to mean the probability of rapid change in the MOC and the consequent impact on climate (affecting temperatures, precipitation, sea level, for example). This project must:

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- \* contribute to the MOC observing system assessment in 2011;
- \* investigate how observations of the MOC can be used to constrain estimates of the probability of rapid MOC change, including magnitude and rate of change;
- \* make sound statistical inferences about the real climate system from model simulations and observations;
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So, it is not unreasonable that the objective of our analysis should be probabilities which are asserted by at least one person (more would be good!).

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- “optimise” the performance of the system



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de Finetti chooses expectation over probability as, if expectation is primitive, then we can choose to make as many or as few expectation statements as we choose, whereas, if probability is primitive, then we must make all of the probability statements before we can make any of the expectation statements, so that we have the option of restricting our attention to whatever subcollection of specifications we are interested in analysing carefully.

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(i) Full Bayes analysis, with complete joint probabilistic specification of all of the uncertain quantities in the problem

or

(ii) Bayes linear analysis, based on a prior specification of the means, variances and covariances of all quantities of interest, where we make expectation, rather than probability, the primitive for the theory, following de Finetti “Theory of Probability”(1974,1975).

de Finetti chooses expectation over probability as, if expectation is primitive, then we can choose to make as many or as few expectation statements as we choose, whereas, if probability is primitive, then we must make all of the probability statements before we can make any of the expectation statements, so that we have the option of restricting our attention to whatever subcollection of specifications we are interested in analysing carefully.

Full Bayes analysis can be more informative if done extremely carefully, both in terms of the prior specification and the analysis. Bayes linear analysis is partial but easier, faster, more robust particularly for history matching and forecasting.

## Bayes linear approach

For very large scale problems a full Bayes analysis is very hard because

- (i) it is difficult to give a meaningful full prior probability specification over high dimensional spaces;
- (ii) the computations, for learning from data (observations and computer runs), particularly when choosing informative runs, may be technically difficult;
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For a full account, see

Michael Goldstein and David Wooff (2007) Bayes Linear Statistics: Theory and Methods, Wiley.



## Bayes linear adjustment

Bayes Linear adjustment of the mean and the variance of  $y$  given  $z$  is

$$\begin{aligned} E_z[y] &= E(y) + \text{Cov}(y, z) \text{Var}(z)^{-1} (z - E(z)), \\ \text{Var}_z[y] &= \text{Var}(y) - \text{Cov}(y, z) \text{Var}(z)^{-1} \text{Cov}(z, y) \end{aligned}$$

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The foundation for the approach is an explicit treatment of temporal uncertainty, and the underpinning mathematical structure is the inner product space (not probability space, which is just a special case).

## Function emulation

Uncertainty analysis, for high dimensional problems, is even more challenging if the function  $f(x)$  is expensive, in time and computational resources, to evaluate for any choice of  $x$ . [For example, large climate models.]

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We use the emulator either to provide a full joint probabilistic description of all of the function values (full Bayes) or to assess expectations variances and covariances for pairs of function values (Bayes linear).

## Form of the emulator

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$$f_i(x) = \sum_j \beta_{ij} g_{ij}(x) \oplus u_i(x)$$



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where  $B = \{\beta_{ij}\}$  are unknown scalars,  $g_{ij}$  are known deterministic functions of  $x$ ,  $u_i(x)$  is a weakly second order stationary stochastic process, with (for example) correlation function

$$\text{Corr}(u_i(x), u_i(x')) = \exp(-(\frac{\|x-x'\|}{\theta_i})^2)$$

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We need careful (multi-output) experimental design to choose informative model evaluations, and detailed diagnostics to check emulator validity.

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Then a relatively small number of evaluations of  $f_i(x)$ , using relations such as

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## Illustration from RAPID (thanks to Danny Williamson)

One of the main aims of the RAPIT programme is to assess the risk of shutdown of the AMOC (Atlantic Meridional Overturning Circulation) which transports heat from the tropics to Northern Europe and how this risk depends on the future emissions scenario for CO<sub>2</sub>.



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We had access to some runs of FAMOUS (a lower resolution model), which consisted of 6 scenarios for future CO<sub>2</sub> forcing, and between 40 and 80 runs of FAMOUS under each scenario, with different parameter choices.

[And very little time to do the analysis.]

## Design

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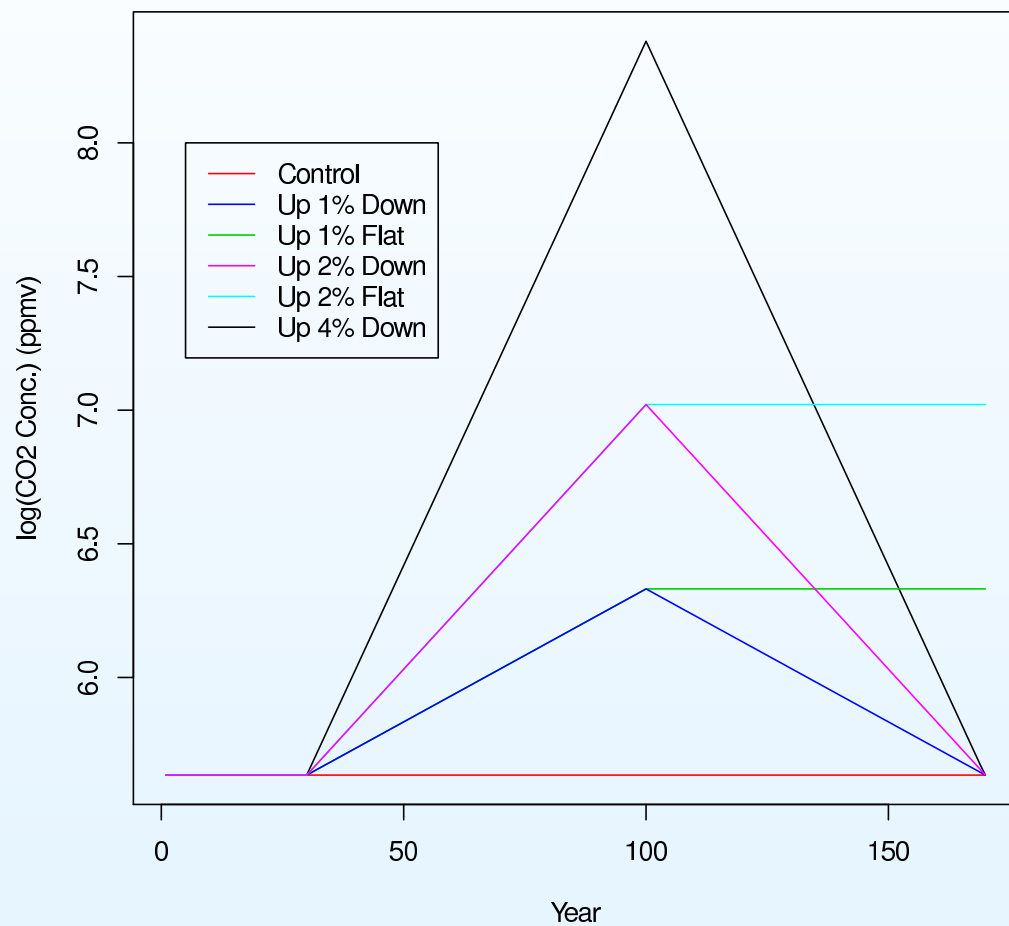
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Our output of interest was a 170 year time series of AMOC values. The series is noisy and the location and direction of spikes in the series was not important. Interest concerned aspects such as the value and location of the smoothed minimum of the series and the amount that AMOC responds to CO<sub>2</sub> forcing and recovers if CO<sub>2</sub> forcing is reduced.

# CO2 Scenarios

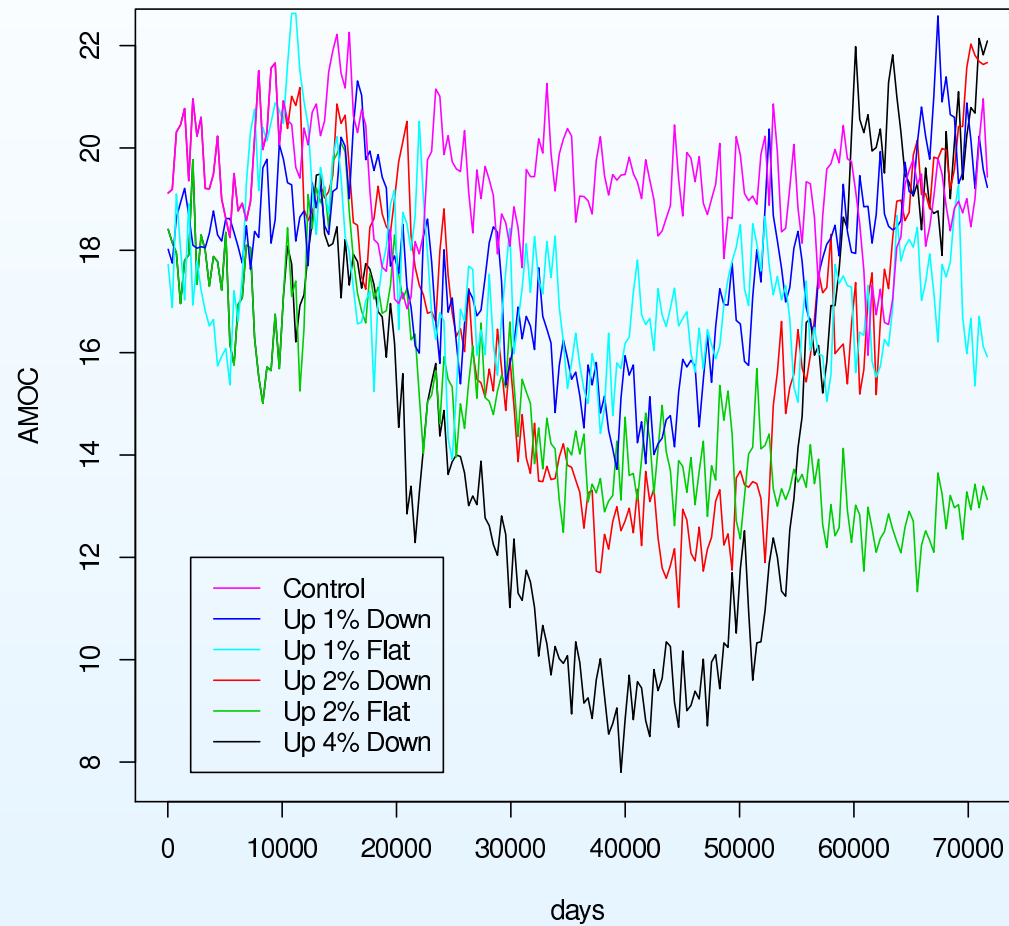
Log CO2 concentration trajectories





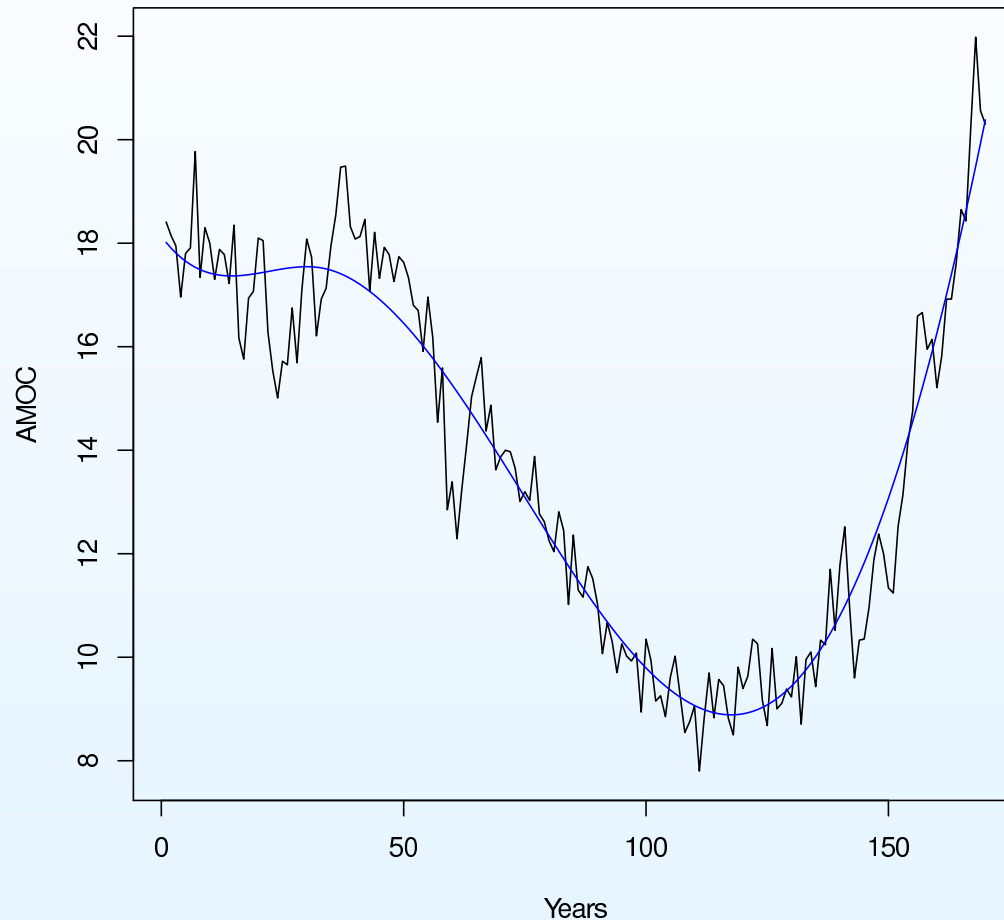
# Famous Scenarios

FAMOUS AMOC Scenarios



# Smoothing

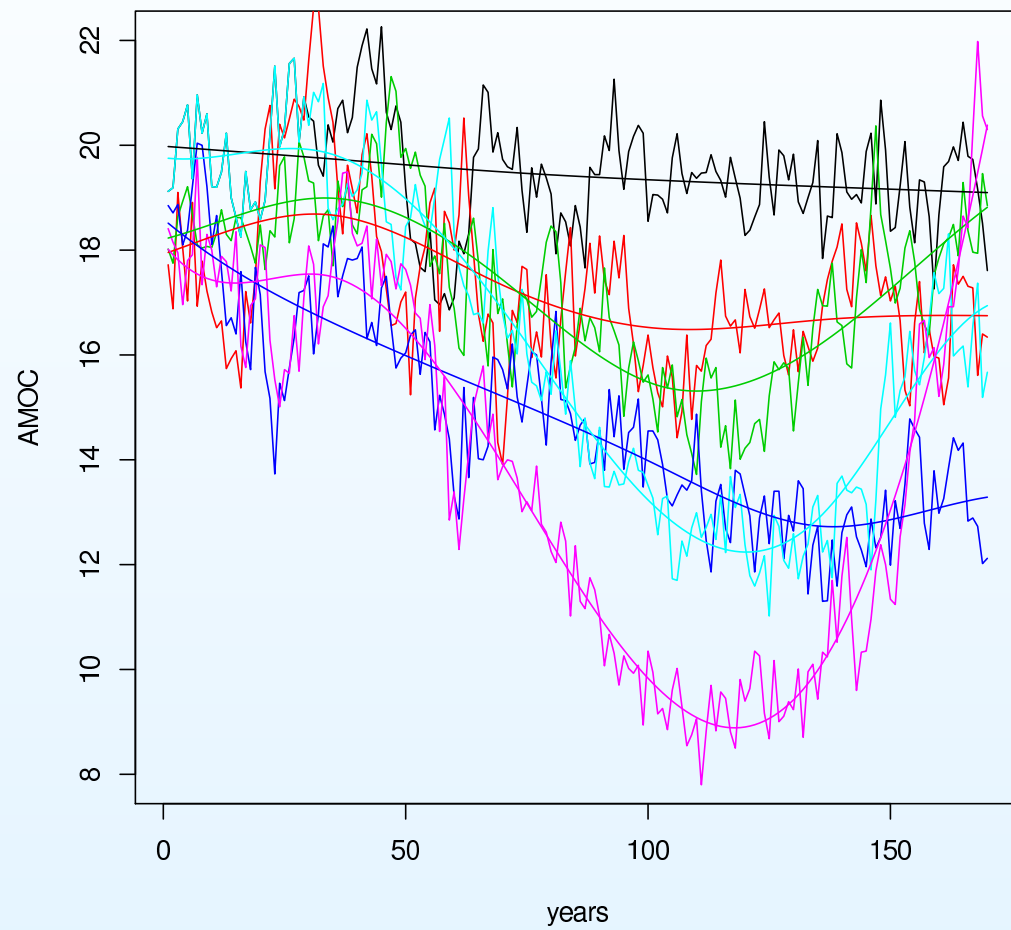
AMOC Up 4% down



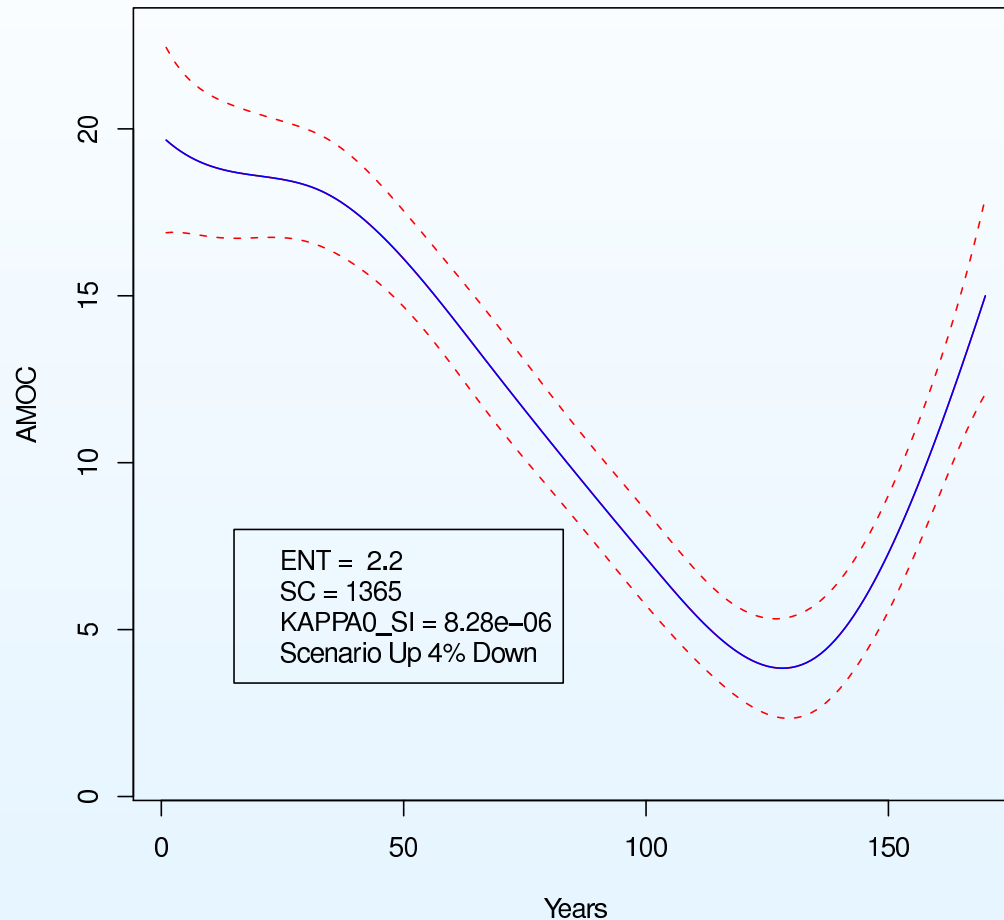
We smooth by fitting splines  $f^s(x, t) = \sum_j c_j(x) B_j(t)$  where  $B_j(t)$  are basis functions over  $t$  and  $c_j(x)$  are chosen to give the 'best' smooth fit to the time series.

# Smoothing

Splines for each scenario



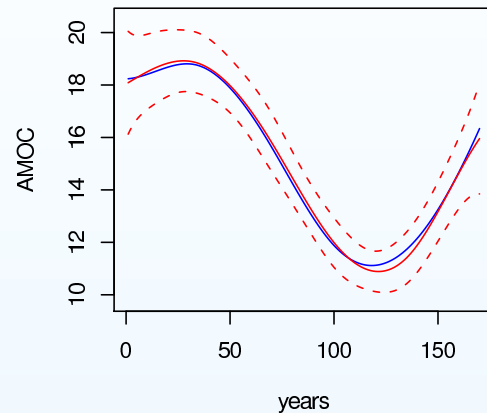
FAMOUS Emulator



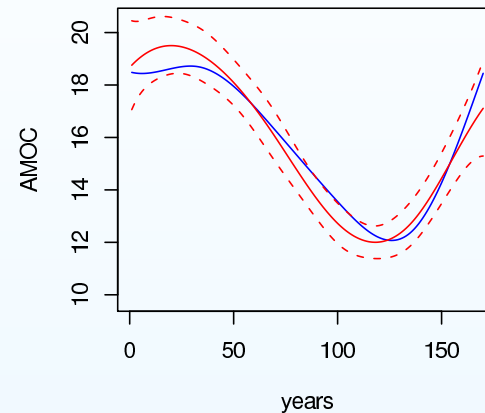
We emulate  $f^s$  by emulating each coefficient  $c_j(x)$  in  
 $f^s(x, t) = \sum_j c_j(x) B_j(t)$  (separately for each CO2 scenario)

## Diagnostics (leave one out)

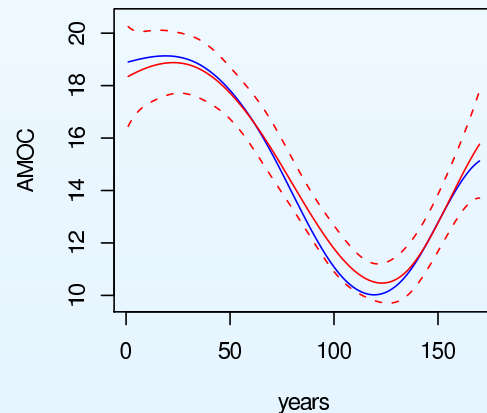
LOO plot for data point 2



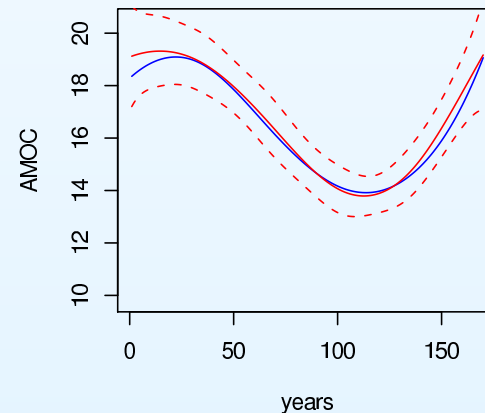
LOO plot for data point 6



LOO plot for data point 7



LOO plot for data point 9



We test our approach by building emulators leaving out each observed run in turn, and checking whether the run falls within the stated uncertainty limits.

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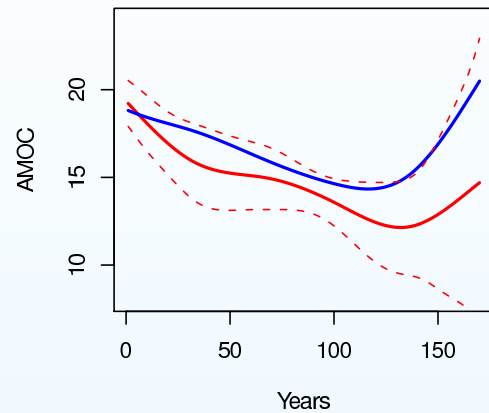
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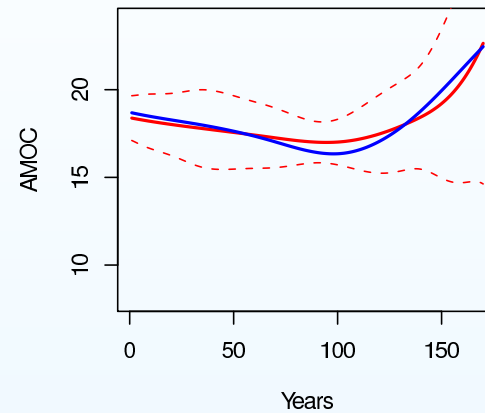
[4] Diagnostic checking, tuning etc.

# Emulating HadCM3:diagnostics

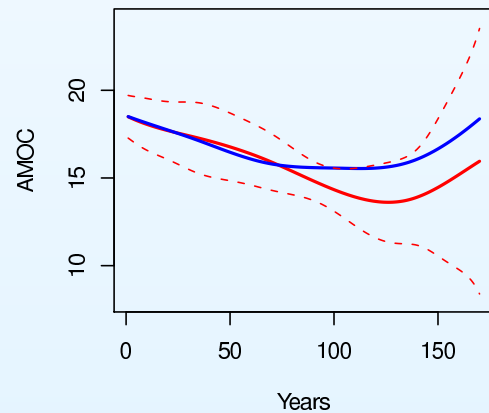
LOO plot for data point 1



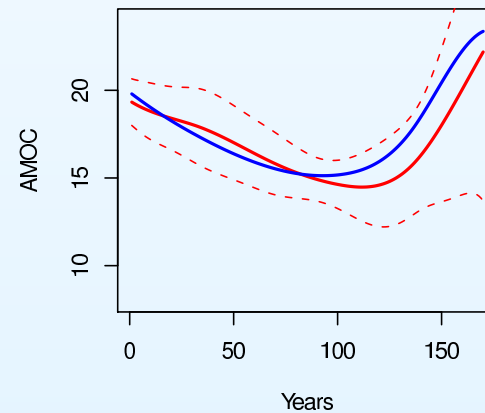
LOO plot for data point 3



LOO plot for data point 10

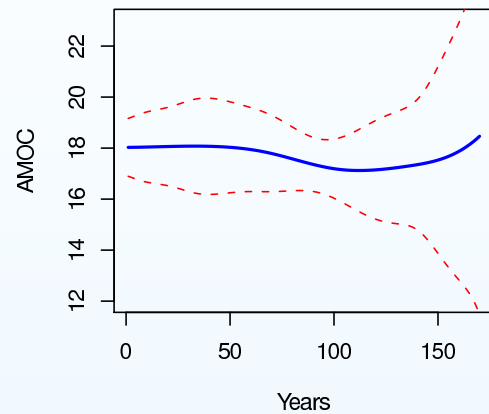


LOO plot for data point 15

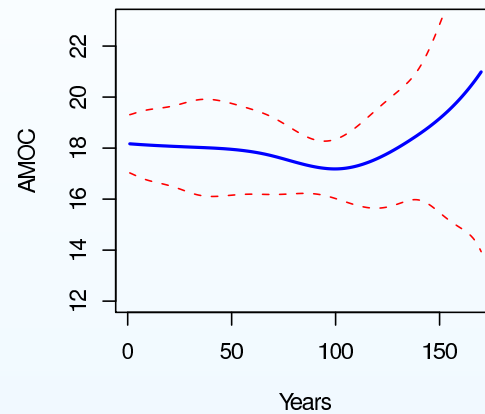


# Emulating HadCM3

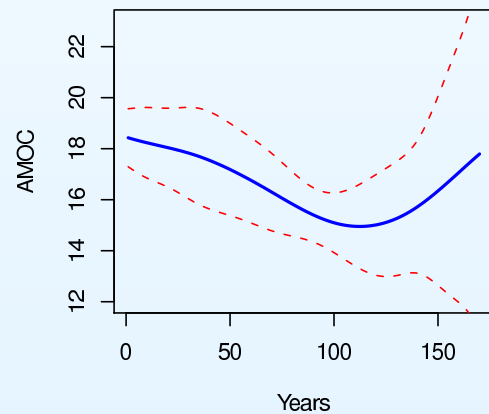
HadCM3 Emulator Up 0.5% Down 0%



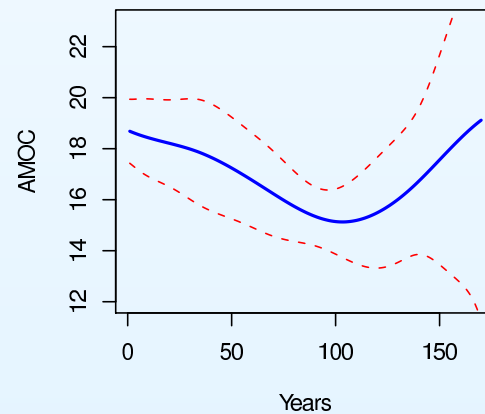
HadCM3 Emulator Up 0.5% Down 0.5%



HadCM3 Emulator Up 1.5% Down 0.5%



HadCM3 Emulator Up 1.5% Down 1%



## Example: Oil Reservoirs

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- Example - Oil field containing 650 wells, 1 million plus grid cells (permeability, porosity, fault lines, etc.). Previous history match took one man-year of effort. Our methods found a match using 32 runs, each lasting 4 hours and automatically chosen with a overall fourfold improvement in fit.



## Inputs and outputs

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Each cell in the reservoir has a collection of associated input parameters, such as permeability and porosity. There are also other parameters, such as Fault transmissibility, Aquifer features, Saturation properties

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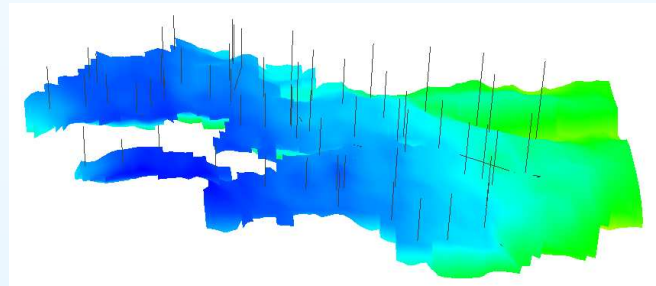
- The model outputs comprise the behaviour of the various wells and injectors in the reservoir
- Output is a time series on the following variables for each well
  - *Pressures* Bottom-hole pressure, Tubing head pressure
  - *Production/Injection rates and totals* for each of oil, water and gas.
  - *Fluid ratios* Water cut, Gas-oil ratio
- The resolution of the time series can be varied from months to years
- With a large number of wells, daily output, or a long operating period there will be a *lot* of output data

## A reservoir example: (thanks to Jonathan Cumming)

The model, based on grid size  $38 \times 87 \times 25$ , with 43 production and 13 injection wells, simulates 10 years of production, 1.5–3 hours per simulation.

**Inputs** Field multipliers for porosity ( $\phi$ ), permeabilities ( $k_x, k_z$ ), critical saturation ( $crw$ ), and aquifer properties ( $A_p, A_h$ )

**Outputs** Oil production rate for a 3-year period, for the 10 production wells active in that period. 4-month averages over the time series



Emulator for reservoir simulator is:  $f_i(\mathbf{x}) = \mathbf{g}_i(\mathbf{x}_{[i]})\beta_i + u_i(\mathbf{x}_{[i]}) + v_i(\mathbf{x})$

$\mathbf{g}_i(\mathbf{x}_{[i]})^T \beta_i$  – a global trend function which captures the gross features,

$\mathbf{x}_{[i]}$  – a subset of inputs which account for most of the variation in  $F$ , the *active variables*,  $u_i(\mathbf{x}_{[i]})$  – a correlated residual process representing the local behaviour in the active variables,  $v_i(\mathbf{x})$  – an uncorrelated ‘nugget’ residual.

## Coarse and Accurate Emulators

The computer model is expensive to evaluate, so we use 'coarse' model,  $F^c$ , to capture qualitative features of  $F$ .  $F^c$  is substantially faster, allowing many model runs. We construct emulator  $f^c$  of  $F^c$  from these runs and a framework linking  $f^c$  and  $f$ . We make (small) number of runs of  $F$ , and update our emulator  $f$ .

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Careful choice of small design to evaluate for full simulator allows us to (Bayes linear) update emulator for  $F$  based on prior emulator and additional runs.

## Emulation Summaries

Well	Time	$\mathbf{x}_{[i]}$	No. Model Terms	Coarse Simulator $R^2$	Accurate Simulator $\tilde{R}^2$
B2	4	$\phi, crw, A_p$	9	0.886	0.951
B2	8	$\phi, crw, A_p$	7	0.959	0.958
B2	12	$\phi, crw, A_p$	10	0.978	0.995
B2	16	$\phi, crw, k_z$	7	0.970	0.995
B2	20	$\phi, crw, k_x$	11	0.967	0.986
B2	24	$\phi, crw, k_x$	10	0.970	0.970
B2	28	$\phi, crw, k_x$	10	0.975	0.981
B2	32	$\phi, crw, k_x$	11	0.980	0.951
B2	36	$\phi, crw, k_x$	11	0.983	0.967

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A conceptually simple alternative is “history matching”, i.e. finding the collection of all input choices  $x$  for which you judge the match of the model to the data,

$\|z - f_h(x)\|$  to be acceptably small, using some “implausibility measure”

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$I(x)$  based on a natural probabilistic metric, accounting for emulator uncertainty, condition uncertain, structural discrepancy, observational error etc. In practice, we proceed by sequentially ruling out regions of  $x$  space which are unlikely to give rise to observed history  $z$ .



## History matching via Implausibility

Using the emulator we can obtain, for each set of inputs  $x$ , the mean and variance,  $E(F_h(x))$  and  $\text{Var}(F_h(x))$ .

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if  $x = x^*$ , then

$$\text{Var}(z_i - E(F_i(x))) = \text{Var}(F_i(x)) + \text{Var}(\epsilon_i) + \text{Var}(e_i).$$

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$$\text{Var}(z_i - E(F_i(x))) = \text{Var}(F_i(x)) + \text{Var}(\epsilon_i) + \text{Var}(e_i).$$

We can therefore calculate, for each output  $F_i(x)$ , the “implausibility” if we consider the value  $x$  to be the best choice  $x^*$ , which is the standardised distance between  $z_i$  and  $E(F_i(x))$ , which is

$$I_{(i)}(x) = |z_i - E(F_i(x))|^2 / [\text{Var}(F_i(x)) + \text{Var}(\epsilon_i) + \text{Var}(e_i)]$$

[Large values of  $I_{(i)}(x)$  suggest that it is implausible that  $x = x^*$ .]

## Using Implausibility measures

The implausibility calculation can be performed univariately, or by multivariate calculation over sub-vectors. The implausibilities are then combined, such as by using  $I_M(x) = \max_i I_{(i)}(x)$ , and can then be used to identify regions of  $x$  with large  $I_M(x)$  as implausible, i.e. unlikely to be good choices for  $x^*$ .

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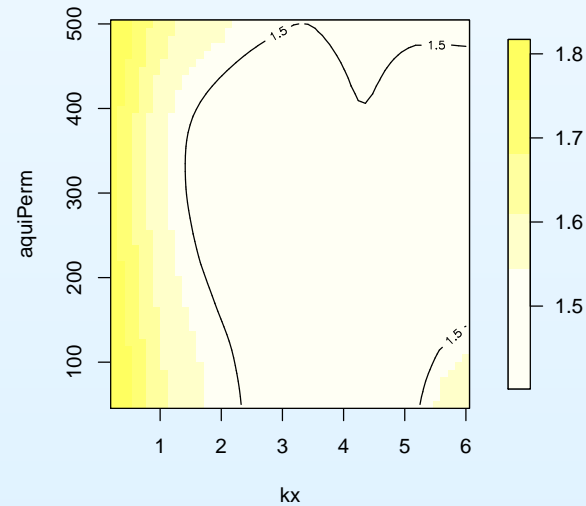
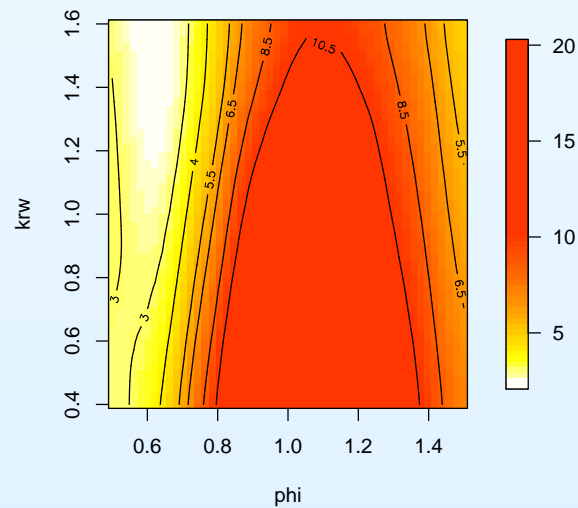
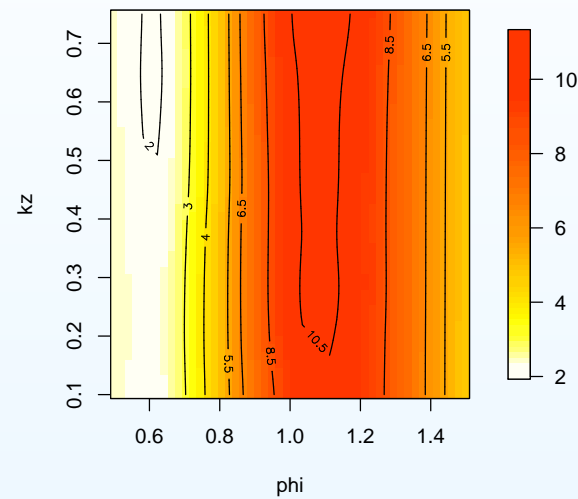
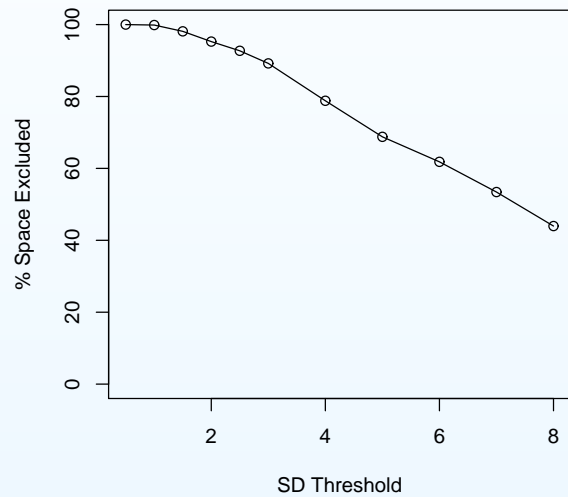
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**Comment:** Even if calibrating, it is good practice to history match first, to check model and (massively) reduce search space.

# Implausibility Results



## Refocusing

- Make the restriction  $\mathcal{X}^* = \{\mathbf{x} : \mathcal{I}(\mathbf{x}) \leq 4\} \simeq \{\mathbf{x} : \phi < 0.79\}$  and eliminate 90% of the input space

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- Now consider final 4 time points in original data, plus an additional point 1 year beyond the end of the previous series to be forecast
- Since reducing the space many of the old model runs are no longer valid, so supplement with additional evaluations
- 262+100 coarse runs, 6+20 accurate runs
- Re-fit the coarse and fine emulators, using the old emulator structure as a starting point

## Forecasting

The mean and variance of  $F(x)$  are obtained from the mean function and variance function of the emulator  $f$  for  $F$ . Using these values, we compute the mean and variance of  $F^* = F(x^*)$  by first conditioning on  $x^*$  and then integrating out  $x^*$ .

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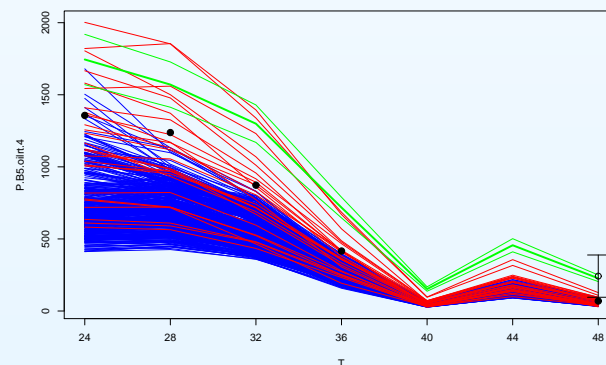
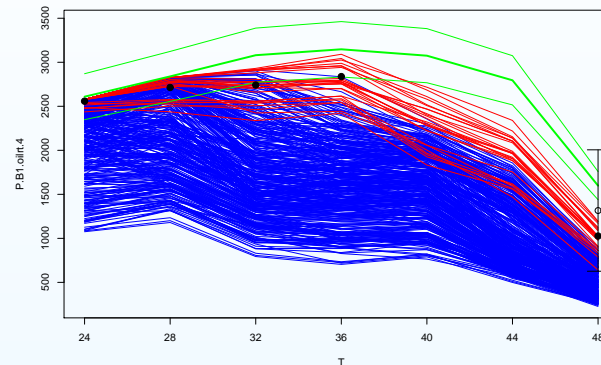
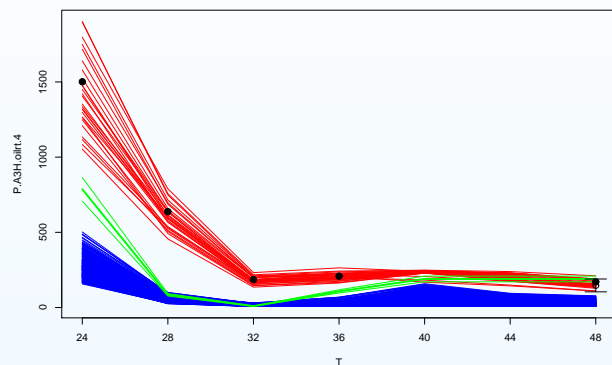
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**Comment** Our computer experiments to forecast  $y_p$  split into two stages

(i) preliminary simulator evaluations to identify the form of emulator, estimate coefficient matrices and refocus

(ii) further simulator evaluations chosen to minimise adjusted forecast variance.

# Forecasting Results



Simulator outputs, observational data and forecasts for each well.

Green lines indicate  $z$  with error bounds of  $2sd(e)$ .

Red and blue lines represent the range of the runs of  $F(x)$  and  $F^c(x)$

Solid black dots correspond to  $E(F^*)$ .

The forecast is indicated by a hollow circle with attached error bars.

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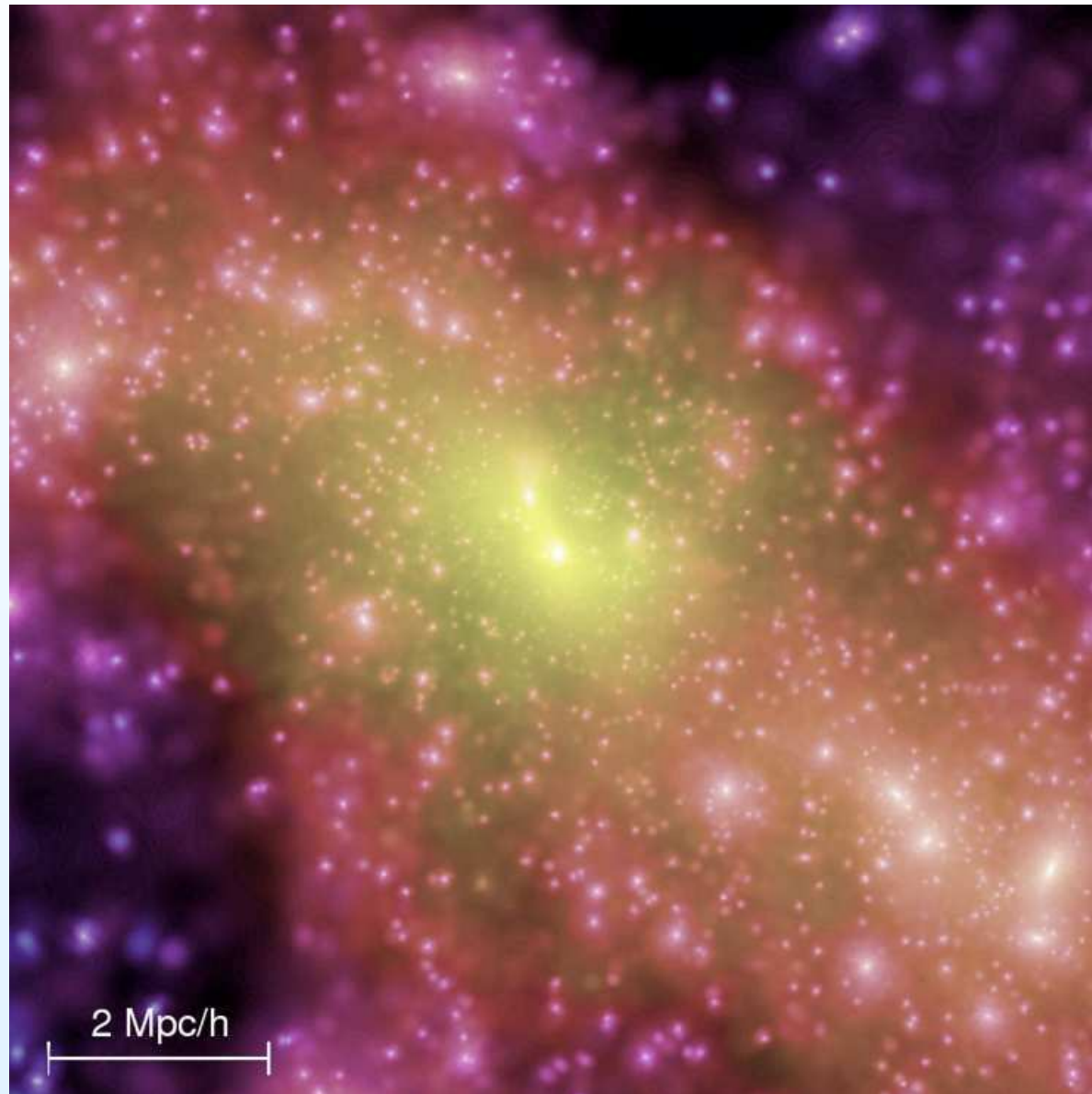
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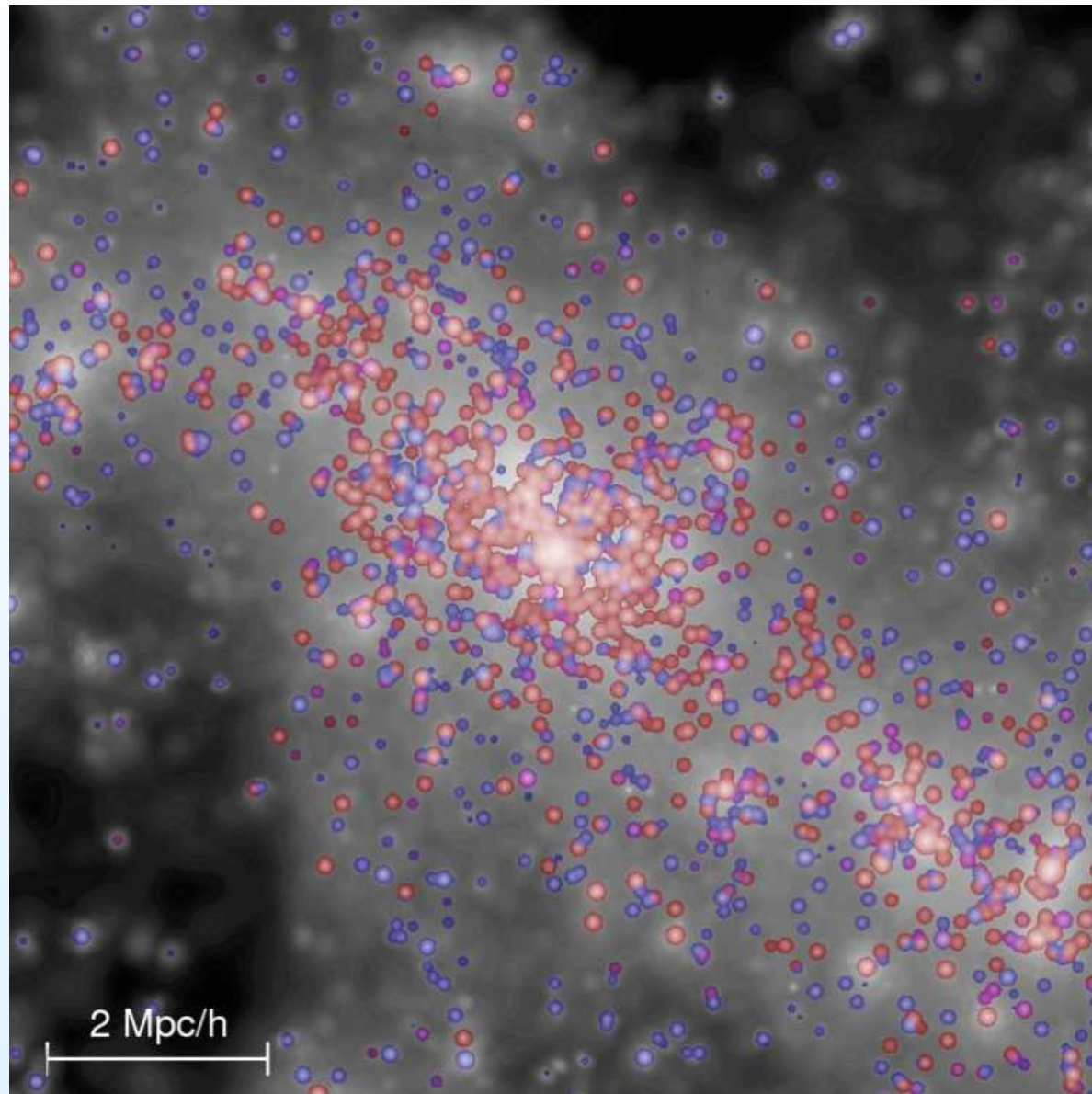
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- It takes approximately 1 day to complete 1 run (using a single processor).
- The Galform model produces lots of outputs, some of which can be compared to observed data from the real Universe.

# The Dark Matter Simulation



# The Galform Model



## Inputs

To perform one run, we need to specify the following 17 inputs:

<b>vhotdisk:</b>	100 - 550	<b>VCUT:</b>	20 - 50
<b>aReheat:</b>	0.2 - 1.2	<b>ZCUT:</b>	6 - 9
<b>alphacool:</b>	0.2 - 1.2	<b>alphastar:</b>	-3.2 - -0.3
<b>vhotburst:</b>	100 - 550	<b>tau0mrg:</b>	0.8 - 2.7
<b>epsilonStar:</b>	0.001 - 0.1	<b>fellip:</b>	0.1 - 0.35
<b>stabledisk:</b>	0.65 - 0.95	<b>fburst:</b>	0.01 - 0.15
<b>alphahot:</b>	2 - 3.7	<b>FSMBH:</b>	0.001 - 0.01
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Galform provides multiple output data sets. Initially we analyse luminosity functions giving the number of galaxies per unit volume, for each luminosity.

B<sub>j</sub> Luminosity: corresponds to density of young (blue) galaxies

K Luminosity: corresponds to density of old (red) galaxies

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<b>epsilonStar:</b>	0.001 - 0.1	<b>fellip:</b>	0.1 - 0.35
<b>stabledisk:</b>	0.65 - 0.95	<b>fburst:</b>	0.01 - 0.15
<b>alphahot:</b>	2 - 3.7	<b>FSMBH:</b>	0.001 - 0.01
<b>yield:</b>	0.02 - 0.05	<b>eSMBH:</b>	0.004 - 0.05
<b>tdisk:</b>	0 - 1		

Galform provides multiple output data sets. Initially we analyse luminosity functions giving the number of galaxies per unit volume, for each luminosity.

B<sub>j</sub> Luminosity: corresponds to density of young (blue) galaxies

K Luminosity: corresponds to density of old (red) galaxies

We choose 11 outputs that are representative of the Luminosity functions and emulate the functions  $f_i(x)$ .

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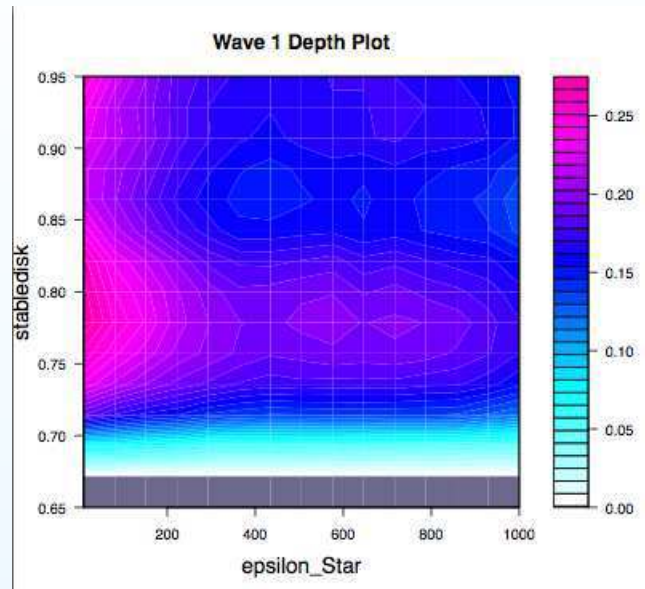
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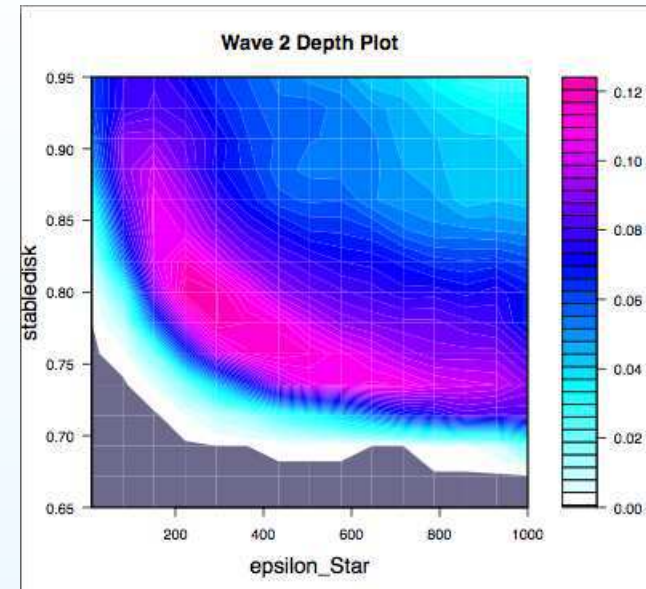
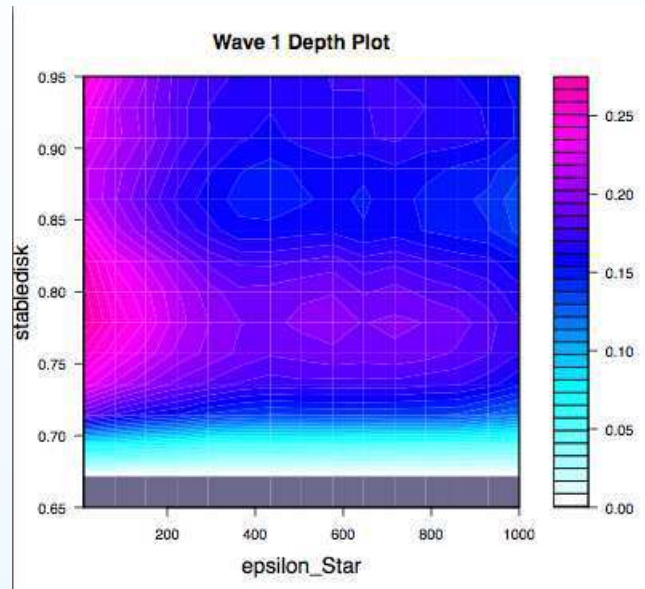
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	No. Model Runs	No. Active Vars	Space Remaining
Wave 1	1000	5	14.9 %
Wave 2	1414	8	5.9 %
Wave 3	1620	8	1.6 %
Wave 4	2011	10	0.12 %

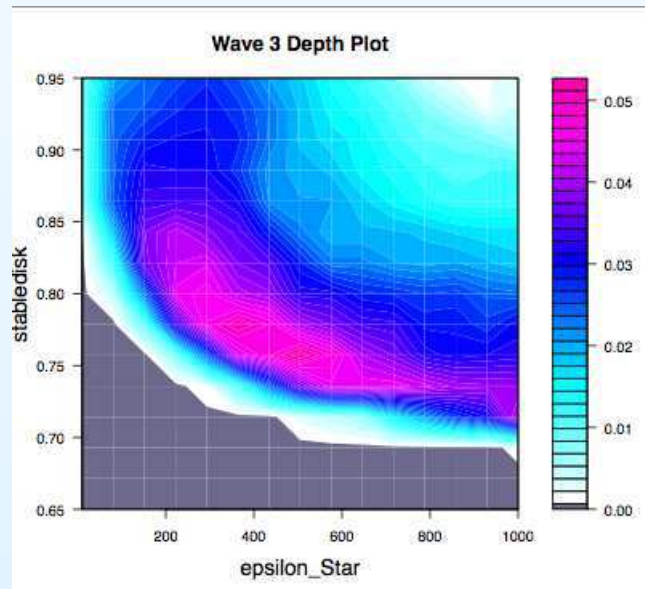
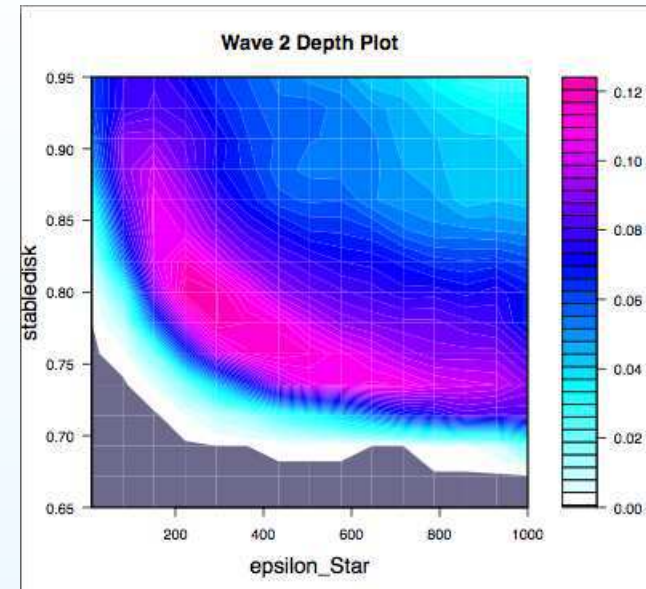
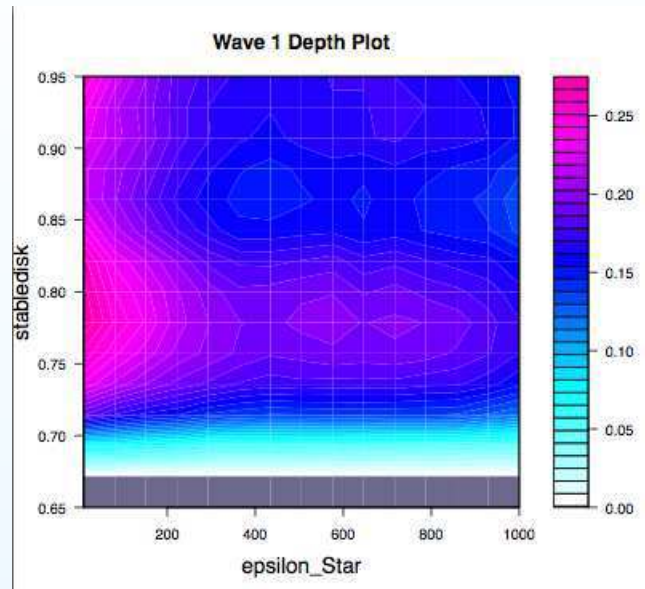
## 2D Implausibility Projections: Wave 1 (14%) to Wave 4 (0.12%)



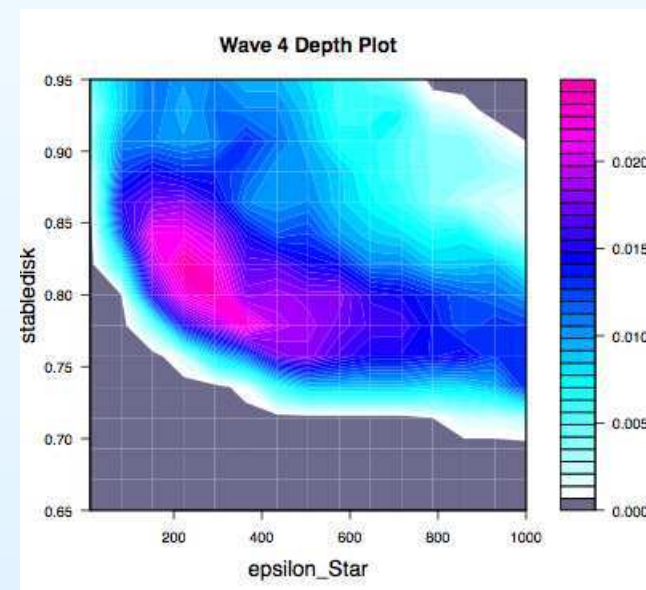
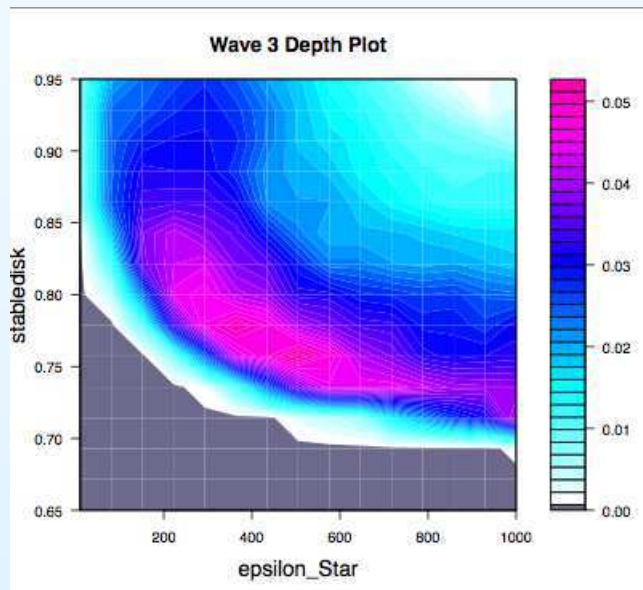
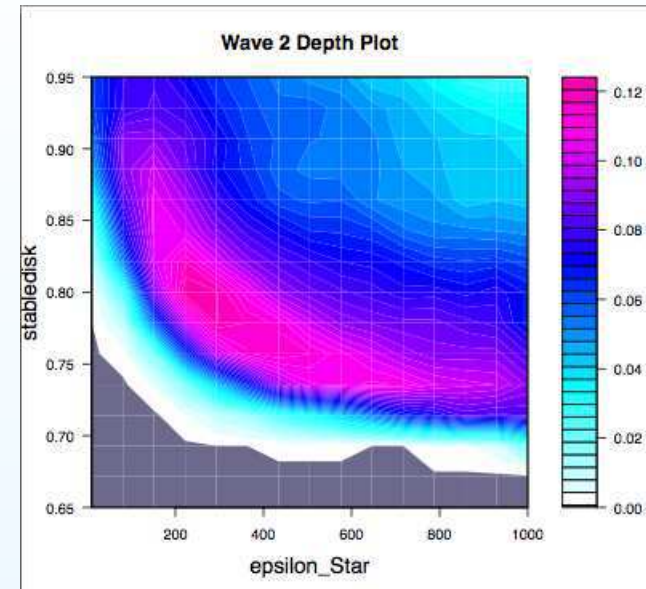
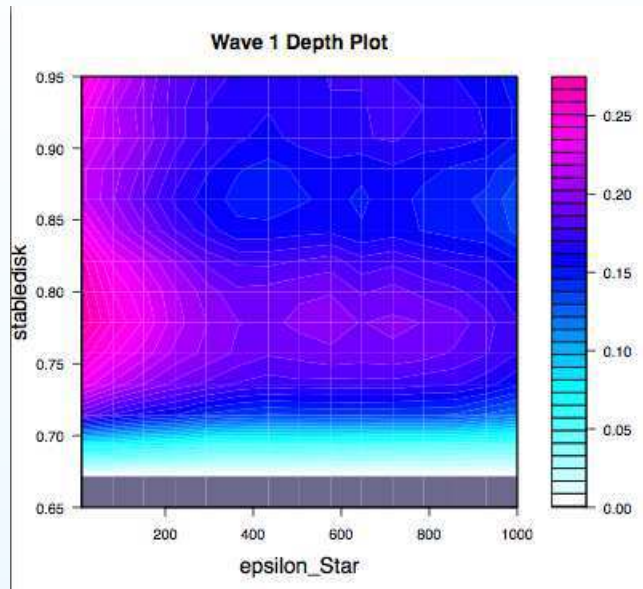
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[A] The difference between our simulator and the reified form.

[B] The difference between the reified form at the physically appropriate choice of  $x$  and the actual system behaviour  $y$ .

## Relating models and the system

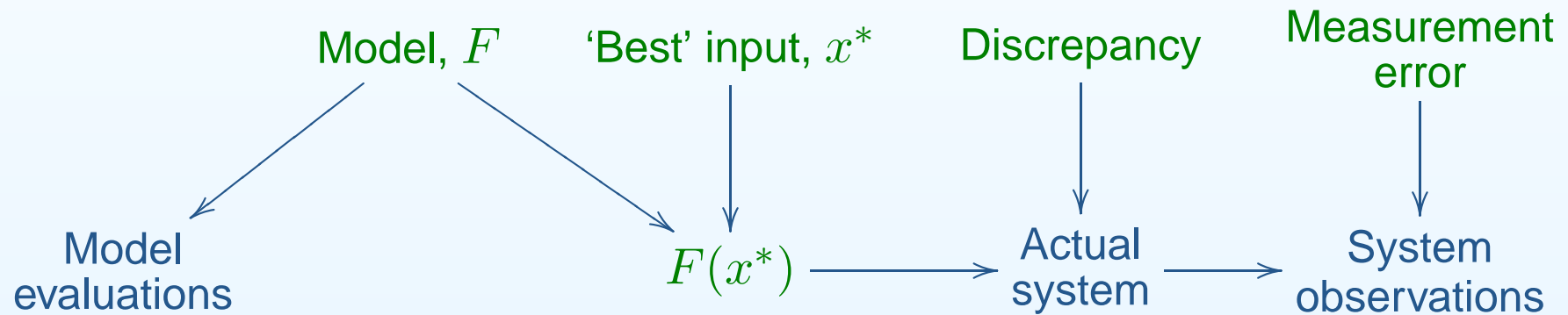
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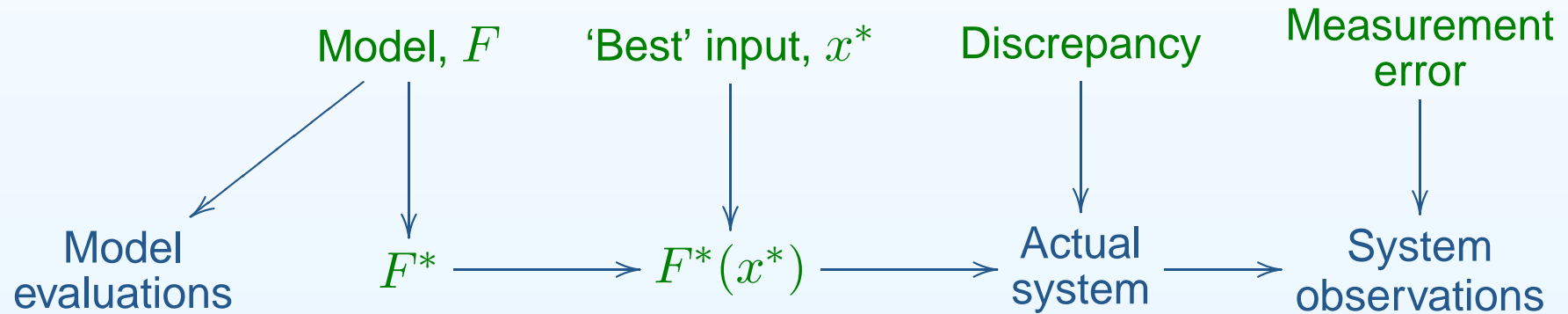
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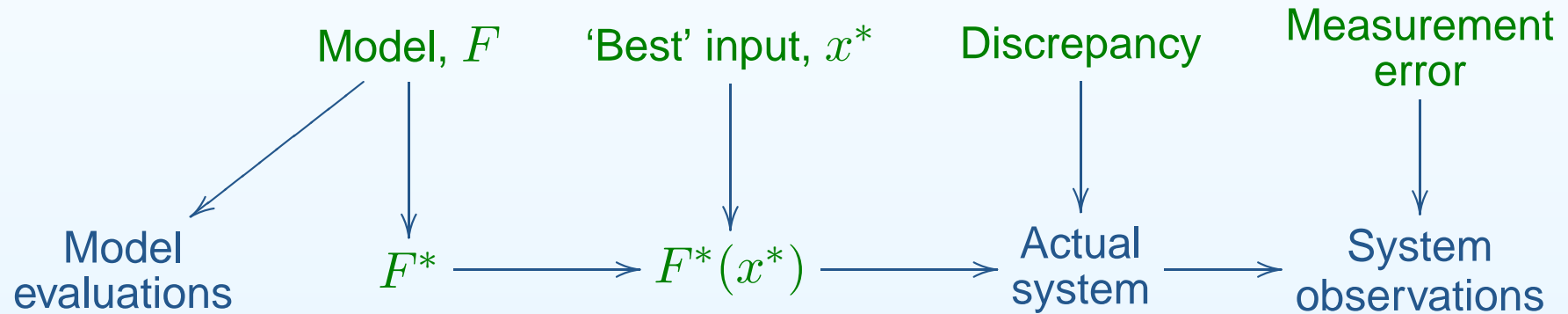
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A collection of simulators  $F_1, F_2, \dots$  is jointly informative for  $y$ , as the simulators are jointly informative for  $F^*$ .



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Structured reification improves on this with systematic modelling for all aspects of model deficiency whose effects we can consider explicitly.

All our calibration and forecasting methodology is unchanged - all that has changed is our description of the joint covariance structure.

## A Reified influence diagram

$$\left[ F_{h:[n]}^1(x), \dots, F_{h:[n]}^m(x) \right]$$

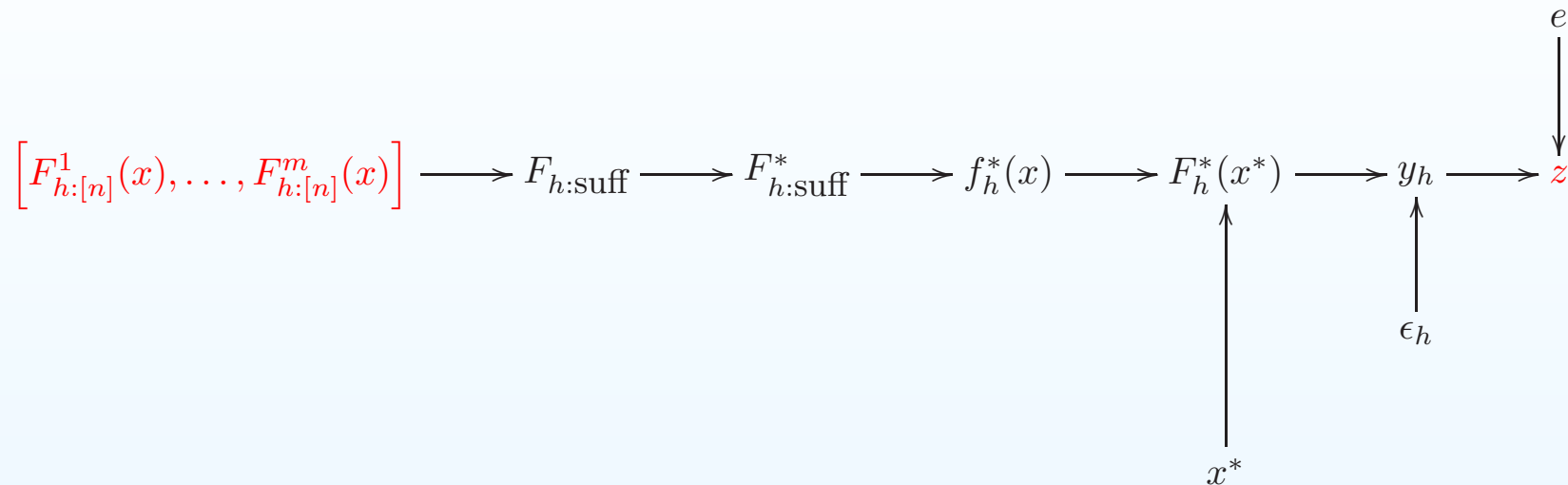
Evaluations of the simulator at each of  $m$  initial conditions  
for historical components of simulator

## A Reified influence diagram

$$\left[ F_{h:[n]}^1(x), \dots, F_{h:[n]}^m(x) \right] \longrightarrow F_{h:\text{suff}} \longrightarrow F_{h:\text{suff}}^* \longrightarrow f_h^*(x)$$

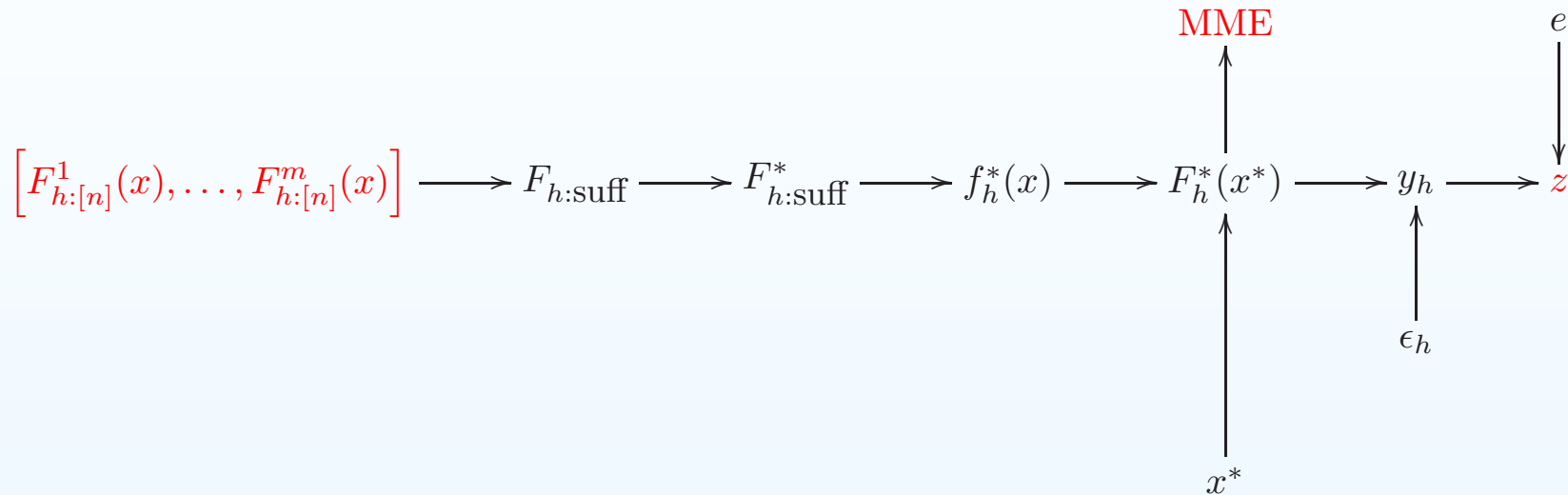
Global information  $F_{h:\text{suff}}$  (from second order exchangeability modelling).  
passes to Reified global form and to reified emulator.

## A Reified influence diagram



Link with  $x^*$  to reified function, at true initial condition, linked to data  $z$

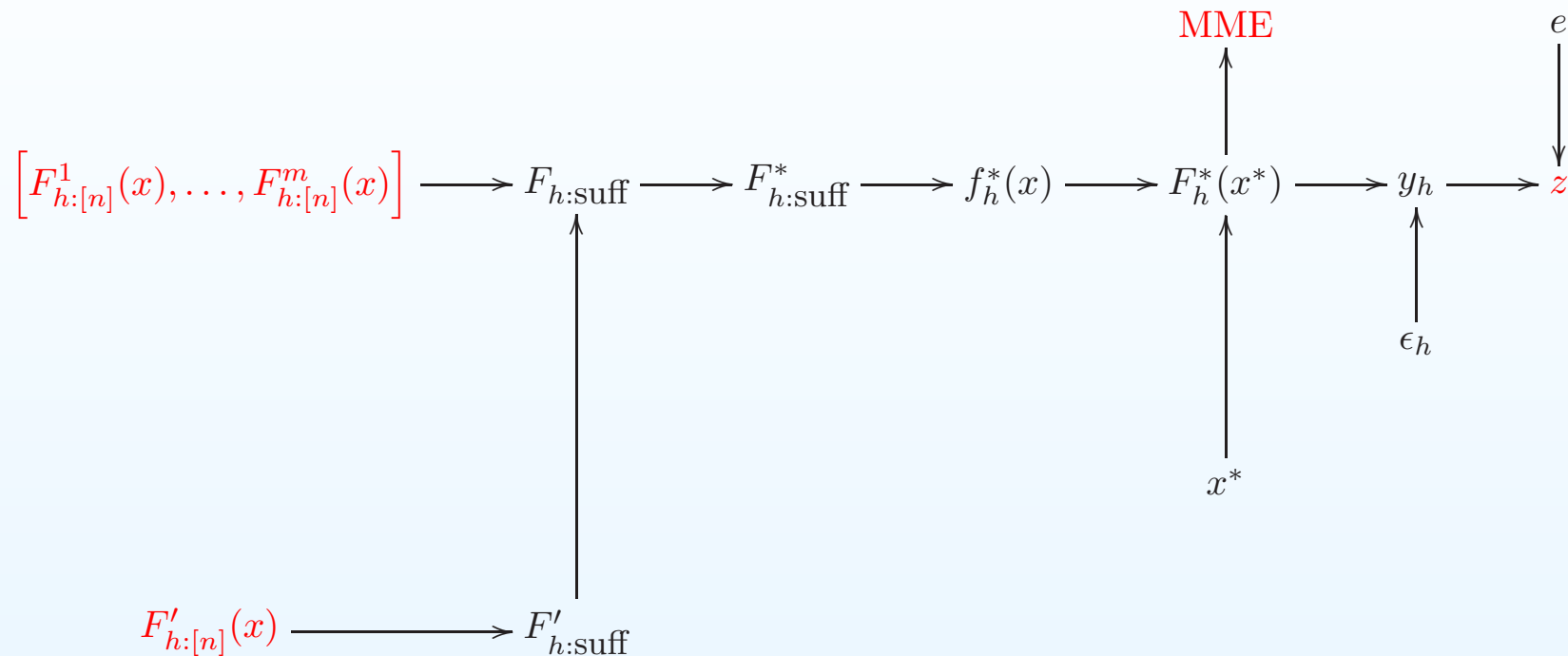
## A Reified influence diagram



Add observation of a related multi-model ensemble (MME) consisting of tuned runs from related models (more exchangeability modelling).

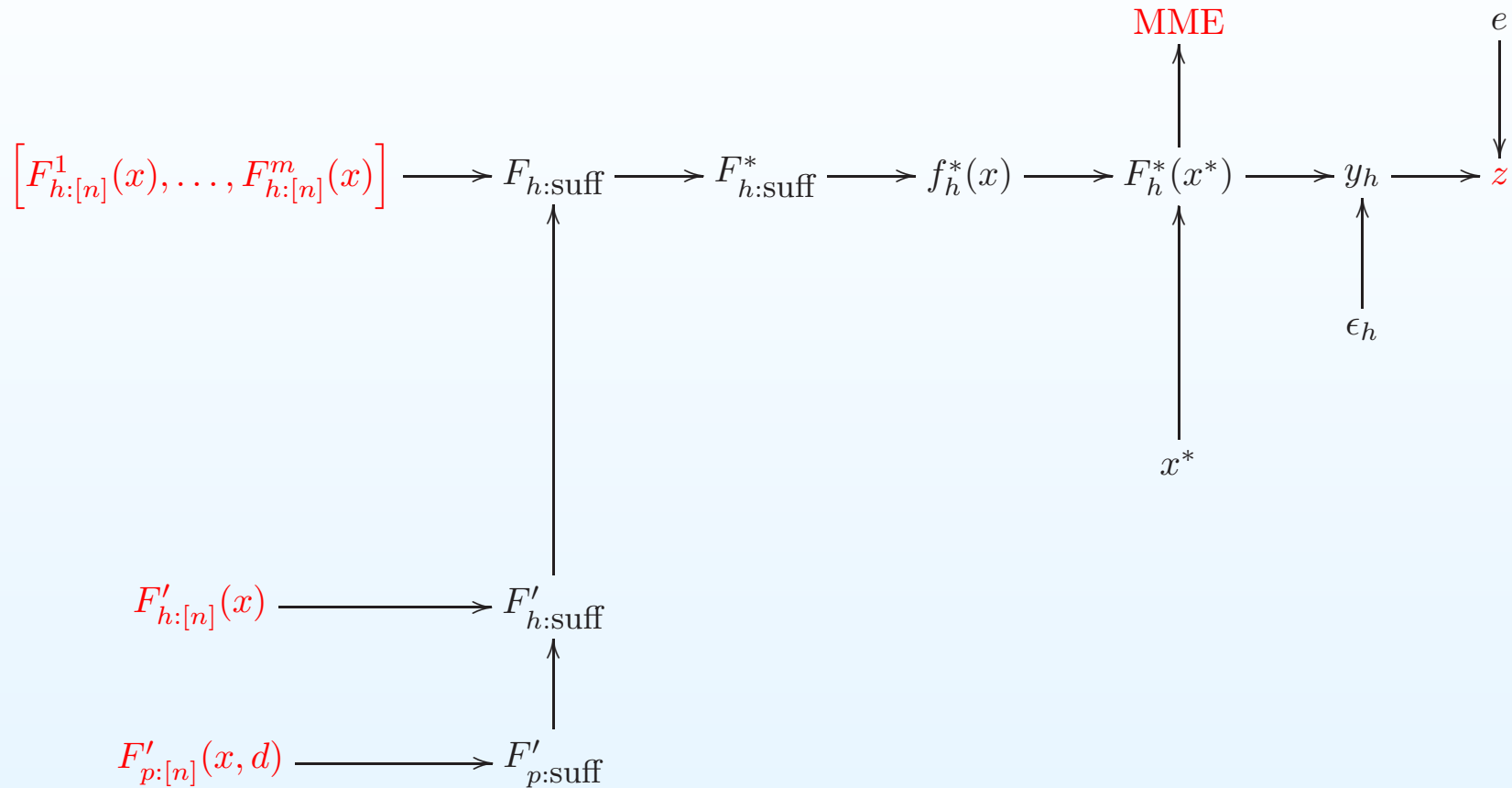


# A Reified influence diagram



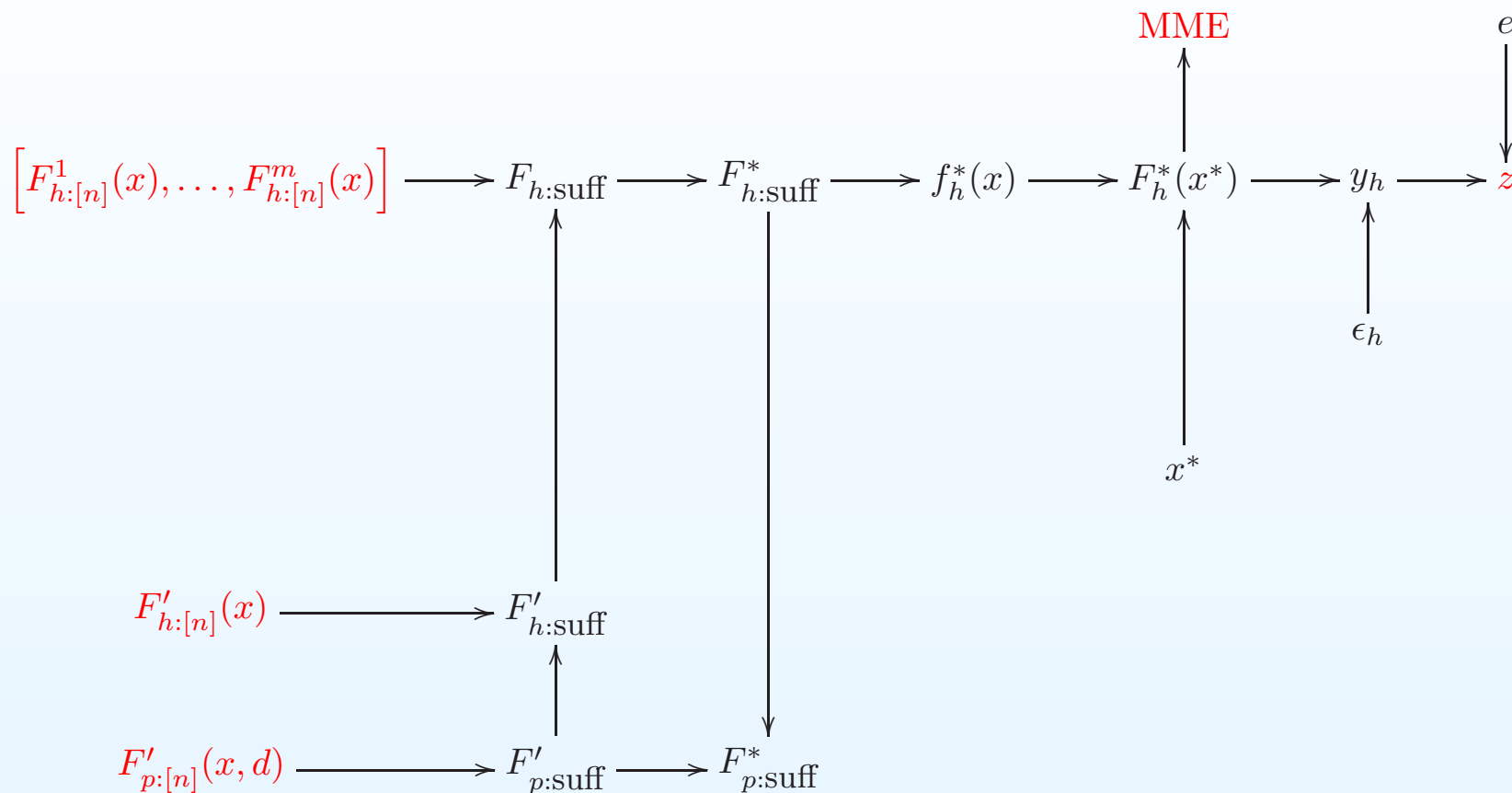
Add a set of evaluations from a fast approximation

# A Reified influence diagram



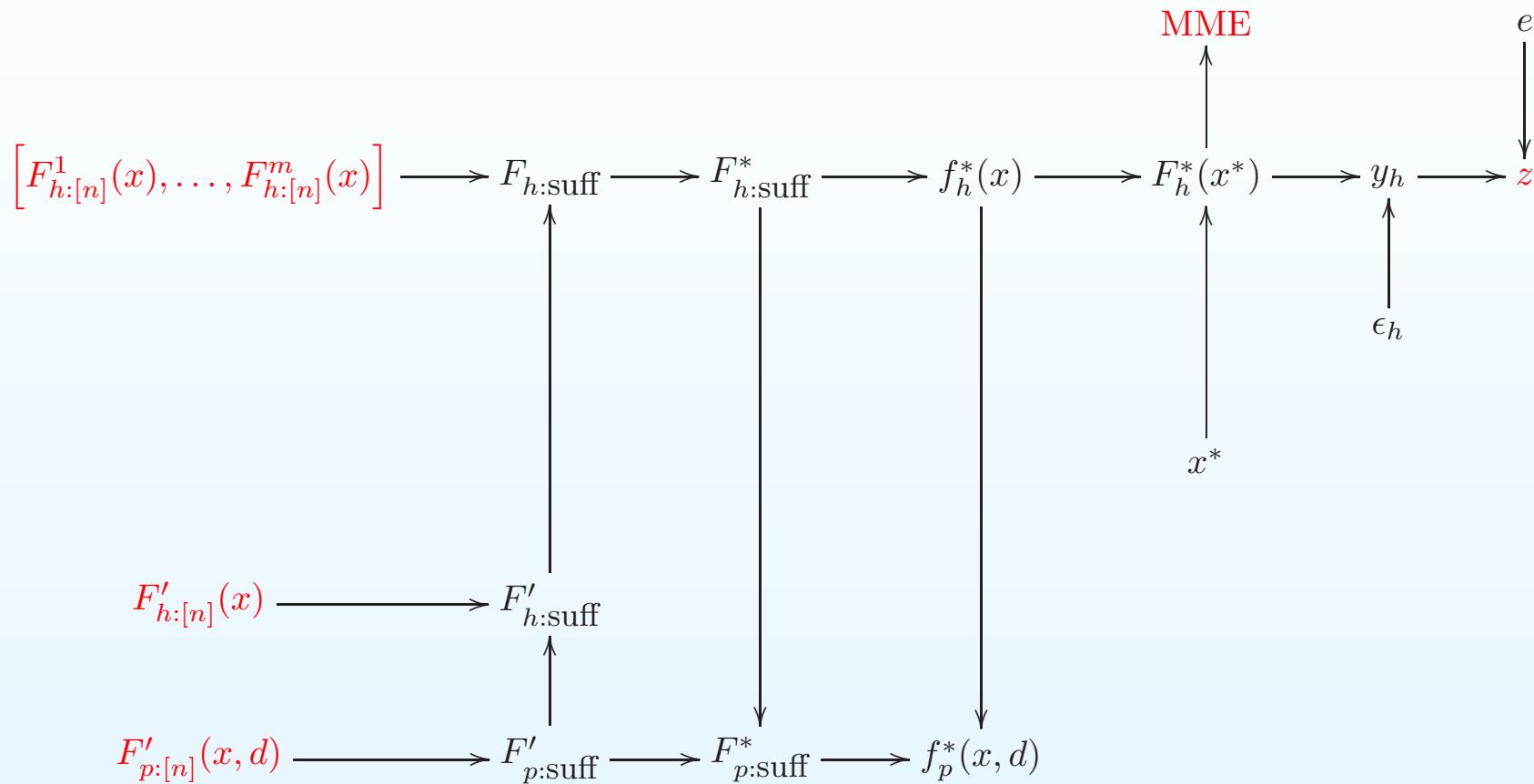
Add evaluations of fast simulator for outcomes to be predicted, with decision choices  $d$

# A Reified influence diagram



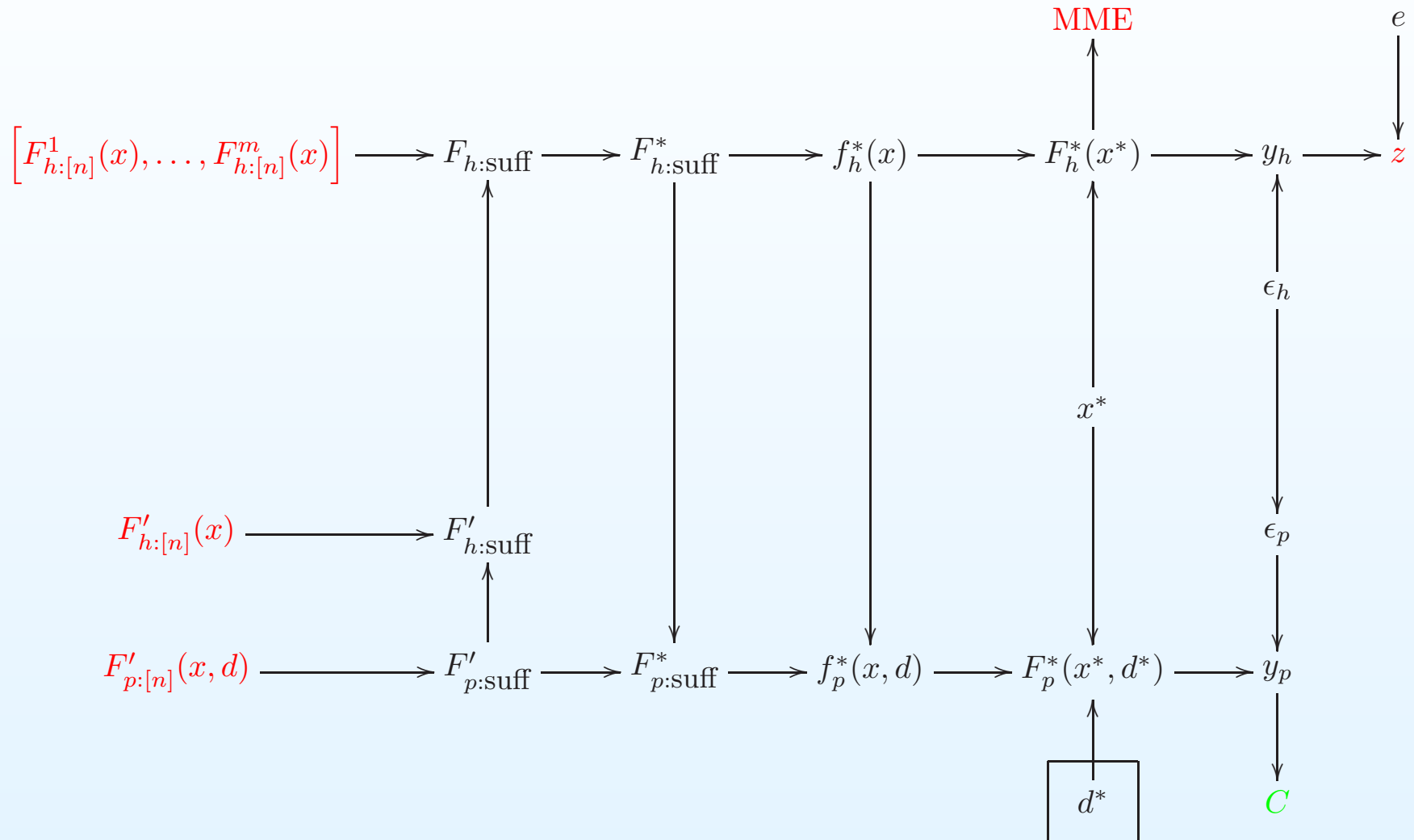
Link to reified global terms for quantities to be predicted

# A Reified influence diagram



And to reified global emulator, based on inputs and decisions

## A Reified influence diagram



And link, through true future values  $y_p$ , to the overall utility cost  $C$  of making decision choice  $d^*$  [Attach more models to diagram at  $F^*(x^*)$ ]

## Concluding comments

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In particular,

Bayesian multivariate, multi-level, multi-model emulation,  
careful structural discrepancy modelling  
and iterative history matching

gives a great first pass treatment for most large modelling problems.

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